

Scintillators with high-energy-resolution and low-intrinsic-activity

F. Quarati^{ab*}, J. van de Biezen^c, P. Dorenbos^a, G. von der Goenna^d, C. Hansson^c, Alan Owens^c, L. Parthier^e, P. Schotanus^f, F. Seifert^e, M. Selle^e, T. Toepfer^d and A. Weisleder^d

^aDelft University of Technology, TNW, Mekelweg 15, 2629JB Delft, The Netherlands

^bPraesepe BV, Lorentzplein 7, 2012 HG Haarlem, The Netherlands

^cEuropean Space Agency, ESTEC, Keplerlaan 1, 2201AZ Noordwijk, The Netherlands

^dHellma Materials GmbH, Moritz-von-Rohr Strasse 1, 07745 Jena, Germany

^eSchott AG, Advanced Materials, Moritz-von-Rohr Strasse 11, 07745 Jena, Germany

^fScionix BV, Regulierenring 5, 3981LA Bunnik, The Netherlands

Lanthanum halides scintillators, namely LaBr_3 and LaCl_3 , have superior scintillation properties and provide the best energy resolution for gamma-ray detection. However, the presence of natural isotope ^{138}La generates an intrinsic activity, well above 1 Bq/cm^3 , which represents a major shortcoming for their application in low count rate experiments. Among other applications, for space applications it is desired the development of low intrinsic activity scintillators which also maintain the spectroscopic performances of LaBr_3 , in terms of energy resolution and detection efficiency. At now two scintillators have been developed by our collaboration with capabilities to fulfill this requirement: CeBr_3 and SrI_2 .

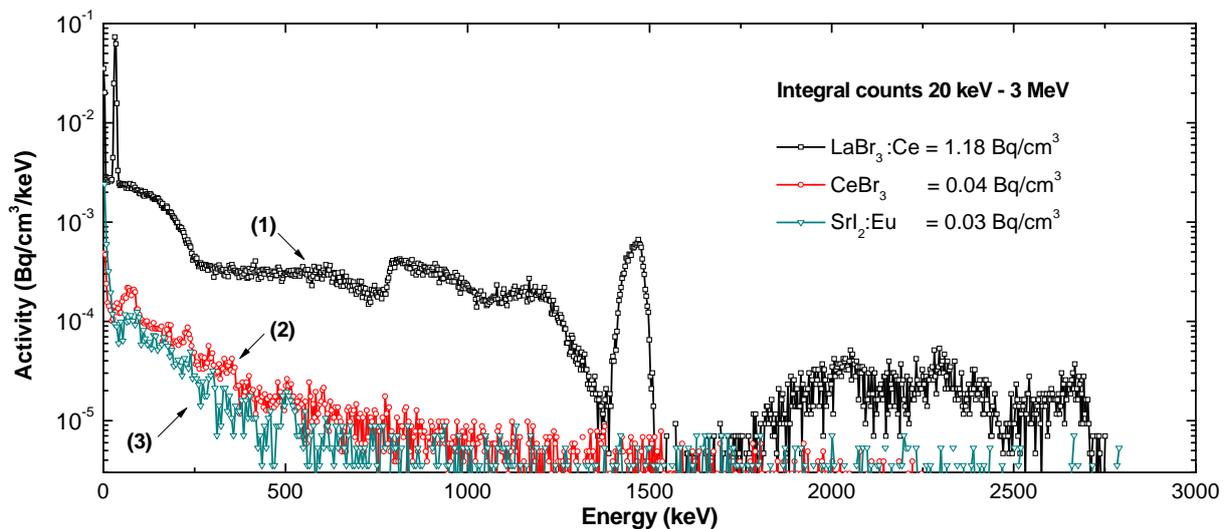


Fig. 1. Comparison of intrinsic activity pulse high spectra collected with $\text{LaBr}_3:\text{Ce}$ (1), CeBr_3 (2) and $\text{SrI}_2:\text{Eu}$ (3).

In this communication we briefly summarize the origin of LaBr_3 intrinsic activity and quantify the limitation it might pose. Then, we show the results we have already obtained with CeBr_3 and SrI_2 . Nowadays CeBr_3 is available in size of $2'' \times 2''$ and larger and does not present any detectable intrinsic activity, neither at level of contaminants. Energy resolution achieved so far is less than 4% at 662 keV and its fundamental limit currently under investigation. At the energy of 1.4 MeV, used for the detection of ^{40}K , which is of interest for gamma-ray remote sensing of planetary surfaces, a CeBr_3 spectrometer is about 8 times more sensitive than LaBr_3 .

* corresponding author; e-mail: f.g.a.quarati@tudelft.nl